Large N Limit of the O(N) Linear Sigma Model via Stochastic Quantization

Rongchan Zhu

Beijing Institute of Technology

Joint work with Hao Shen, Scott Smith and Xiangchan Zhu

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2 Large N limit of the dynamics

Invariant measures



 Recall the free field in the Euclidean quantum field theory. The usual free field on the torus T^d is heuristically described by the following probability measure:

$$\nu(\mathrm{d}\Phi) = C_N^{-1} \Pi_{x \in \mathbb{T}^d} \mathrm{d}\Phi(x) \exp\left(-\int_{\mathbb{T}^d} (|\nabla \Phi|^2 + m\Phi^2) \mathrm{d}x\right),$$

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- The free field describes particles which do not interact.

The Φ_d^4 model is the simplest non-trivial Euclidean quantum field: $C_N^{-1} \Pi_{x \in \mathbb{T}^d} \mathrm{d}\Phi(x) \exp\left(-\int_{\mathbb{T}^d} (|\nabla \Phi(x)|^2 + m\Phi^2(x))\right)$

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Stochastic quantization of Euclidean quantum fields: getting the Φ_d^4 field as stationary distributions (limiting distributions) of stochastic processes, which are solutions to SPDE (see [Parisi,Wu 81], [G. Jona-Lasinio,P. K. Mitter 85], [Albeverio, Röckner 91], [Da Prato, Debussche 03]).

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- Other models

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$$\mathcal{L}\Phi_i = -\frac{1}{N}\sum_{j=1}^N \Phi_j^2 \Phi_i + \xi_i,$$

 $\mathcal{L} = \partial_t - \Delta + m$; $(\xi_i)_{i=1}^N$: independent space-time white noises.

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- Decompose $\Phi_i = Y_i + Z_i$ as Da Prato-Debussche trick for d = 2

$$\begin{aligned} \mathcal{L}Z_i = &\xi_i, \\ \mathcal{L}Y_i = -\frac{1}{N} \sum_{j=1}^{N} (Y_j^2 Y_i + Y_j^2 Z_i + 2Y_j Y_i Z_j \\ &+ 2Y_j \underbrace{Z_i Z_j :}_{Wick \ product} + \underbrace{Z_j^2 :}_{Wick \ product} Y_i + \underbrace{Z_i Z_j^2 :}_{Wick \ product}), \end{aligned}$$

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7:

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- This is not enough since the stopping time may depend on N

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Suppose that d = 2 and (ψ_i, ψ_j) are independent and have the same law and for $p > 1 \mathbf{E} \| \phi_i - \psi_i \|_{C^{-\kappa}}^p \to 0$, as $N \to \infty$.

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• Mean field limit/ Propagation of chaos

$$\begin{split} \Phi_{i} &= Z_{i} + Y_{i}, \ \Psi_{i} = Z_{i} + X_{i} \\ \mathcal{L}Z_{i} &= \xi_{i}, \\ \mathcal{L}Y_{i} &= -\frac{1}{N} \sum_{j=1}^{N} (Y_{j}^{2}Y_{i} + Y_{j}^{2}Z_{i} + 2Y_{j}Z_{j}Y_{i} + 2Y_{j} : Z_{j}Z_{i} : + :Z_{j}^{2} : Y_{i} + :Z_{i}Z_{j}^{2} :), \\ \mathcal{L}X_{i} &= -(\mathbf{E}[X_{j}^{2}]X_{i} + \mathbf{E}[X_{j}^{2}]Z_{i} + 2\mathbf{E}[X_{j}Z_{j}]X_{i} + 2\mathbf{E}[X_{j}Z_{j}]Z_{i}), \\ \text{where } \mu &= \mathbf{E}[\Psi_{i}^{2} - Z_{i}^{2}] = \mathbf{E}[X_{i}^{2}] + 2\mathbf{E}[X_{i}Z_{i}]. \end{split}$$

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Lemma 1

It holds that for $p \geq 2$

$$\frac{1}{N} \mathbf{E} \sup_{t \in [0,T]} \sum_{j=1}^{N} \|Y_{j}\|_{L^{2}}^{2} + \frac{1}{N} \sum_{j=1}^{N} \mathbf{E} \|\nabla Y_{j}\|_{L^{2}(0,T;L^{2})}^{2} + \mathbf{E} \left\|\frac{1}{N} \sum_{i=1}^{N} Y_{i}^{2}\right\|_{L^{2}(0,T;L^{2})}^{2} \lesssim 1,$$

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• The limiting equation

$$(\partial_t - \Delta + m)\Psi_i = \mathcal{L}\Psi_i = -\mu\Psi_i + \xi_i,$$

where $\mu = \mathbf{E}[\Psi_i^2 - Z_i^2], d = 2; \mu = \mathbf{E}[\Psi_i^2], d = 1;$

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• Invariant measure: Gaussian free field

$$\mathcal{N}(0, (m - \Delta)^{-1}), d = 2, 3; \quad \mathcal{N}(0, (m + \mu_0 - \Delta)^{-1}), d = 1,$$

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For d = 1, 2, there exists $m_0 > 0$ such that: for $m \ge m_0$, the Gaussian free field $\mathcal{N}(0, (m - \Delta)^{-1})$ is the unique invariant measure to Ψ .

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Convergence of invariant measure (field)

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$$\nu^{N} = \frac{1}{C_{N}} \exp\left(-2\int_{\mathbb{T}^{d}} \frac{1}{2} \sum_{j=1}^{N} |\nabla \Phi_{j}|^{2} + \frac{m}{2} \sum_{j=1}^{N} \Phi_{j}^{2} + \frac{1}{4N} \left(\sum_{j=1}^{N} \Phi_{j}^{2}\right)^{2} \mathrm{d}x\right) \mathcal{D}\Phi,$$

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- Difficulty: don't know each component of the tight limit is independent; Nonlinear Markov semigroup P^{*}_tν₁ ≠ ∫ P^{*}_tδ_xν₁(dx); not easy to control the nonlinear term as time goes to infinity

Idea of proof: $\mathbb{W}_2(
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• Cancelation from paraprodcts [Gubinelli, Hofmanova 18]

$$\sum_{i=1}^{N} \left[\langle (\Delta - m) Y_i, Y_i \rangle - \frac{1}{N} \sum_{j=1}^{N} \langle 2Y_j \preccurlyeq : Z_i Z_j : +Y_i \preccurlyeq : Z_j^2 :, Y_i \rangle \right]$$

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 $R_N = R_N - \mathbf{E}[R_N] + \mathbf{E}[R_N]$

Observables

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Theorem [Shen, Scott, Zhu, Z. 20/ Shen, Zhu, Z. 21]

Suppose that $\Phi \simeq \nu^N$. For $\kappa > 0$, *m* large enough, the following result holds:

• $\frac{1}{\sqrt{N}} \sum_{i=1}^{N} : \Phi_i^2$: is tight in $B_{2,2}^{-2\kappa}$ for $d = 2 / B_{1,1}^{-1-\kappa}$ for d = 3• $\frac{1}{N} : (\sum_{i=1}^{N} \Phi_i^2)^2$: is tight in $B_{1,1}^{-3\kappa}$ for d = 2

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• Idea: Improved moment estimate for stationary case by independence

$$\mathbf{E}\left[\left(\sum_{i=1}^{N} \|Y_i\|_{L^2}^2\right)^q\right] + \mathbf{E}\left[\left(\sum_{i=1}^{N} \|Y_i\|_{L^2}^2 + 1\right)^q \left(\sum_{i=1}^{N} \|\nabla Y_i\|_{L^2}^2\right)\right] \lesssim 1.$$

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Theorem [Shen, Scott, Zhu, Z. 20/ Shen, Zhu, Z. 21]

Suppose that $\Phi \simeq \nu^N$. For $\kappa > 0$, *m* large enough, the following result holds:

• $\frac{1}{\sqrt{N}} \sum_{i=1}^{N} : \Phi_i^2$: is tight in $B_{2,2}^{-2\kappa}$ for $d = 2 / B_{1,1}^{-1-\kappa}$ for d = 3• $\frac{1}{N} : (\sum_{i=1}^{N} \Phi_i^2)^2$: is tight in $B_{1,1}^{-3\kappa}$ for d = 2

• For *d* = 1, 2,

$$\lim_{N \to \infty} \frac{1}{\sqrt{N}} \sum_{i=1}^{N} : \Phi_i^2 : \neq \lim_{N \to \infty} \frac{1}{\sqrt{N}} \sum_{i=1}^{N} : Z_i^2 :$$
$$\lim_{N \to \infty} \frac{1}{N} : (\sum_{i=1}^{N} \Phi_i^2)^2 : \neq \lim_{N \to \infty} \frac{1}{N} : (\sum_{i=1}^{N} Z_i^2)^2 :$$

• Idea: Improved moment estimate for stationary case by independence

$$\mathbf{E}\left[\left(\sum_{i=1}^{N} \|Y_i\|_{L^2}^2\right)^q\right] + \mathbf{E}\left[\left(\sum_{i=1}^{N} \|Y_i\|_{L^2}^2 + 1\right)^q \left(\sum_{i=1}^{N} \|\nabla Y_i\|_{L^2}^2\right)\right] \lesssim 1.$$

• Integration by parts formula/ Dyson-Schwinger from [Kupiainen 80]

Rongchan Zhu (Beijing Institute of Technology)

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- Other models

Thank you !